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# New K–AR ages of tchabal mbabo alkaline volcano massif, Cameroon volcanic line and adamawa plateau (central Africa)

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#### Abstract

Tchabal Mbabo in Central Africa is a voluminous volcano massif composed of alkali lava series. K-Ar geochronology data obtained from three samples of basanite - trachyandesite composition defined at least two volcanic activities. The first at end Eocene  $(38.22 \pm 0.80 \text{ Ma})$  and the second during Oligocene  $(28.88 \pm 0.61 \text{ and } 28.60 \pm 0.60 \text{ Ma})$ . The distribution of different ages along the Cameroon Volcanic Line is difficult to council with any migration of magmatic activity, as previously suggested. The most realistic scenario for the formation of Cameroon Line is that the ascent of lavas has been favored by the crustal discontinuities inherited from the Pan-African orogeny and reactivated during Mesozoic and Cenozoic. ICP MS and ICP AES analyses show that basanite lavas are the result of 2 % melting of sub-lithospheric mantle source containing garnet and phlogopite phases; the trachyandesites are formed by fractional crystallization of K-feldspar, amphibole, clinopyroxene and Fe-Ti oxides.

Keywords: K-Ar Age; Tchabal Mbabo Volcano; Cameroon Volcanic Line; Adamawa Plateau; Central Africa.

# 1. Introduction

The Tchabal Mbabo massif is a huge Cenozoic volcano that crops out at the cross intersection of well-known magmatic provinces of Cameroon Volcanic Line (CVL) and Adamawa plateau (AP) in Central Africa (Fig. 1). These tectono-magmatic structures have concentrated the major part of Central Africa magmatism with those of Benue trough. They are bordered by the Central Africa Shear Zone (CASZ) which crosscuts both the continental crust and the upper mantle down to a depth of 190 km (Dorbath et al. 1986). The Tchabal Mbabo massif, 2500 m height and stretching 30 x 54 km, marks the Adamawa plateau. It overhangs the western Koutine plain in the West, Kontcha basin to the North, Galim and Sambolabo lowlands to the East and South, respectively. The previous K-Ar (Dunlop 1983; Fitton& Dunlop 1985;Fitton 1987) and <sup>40</sup>Ar/<sup>39</sup>Ar ages (Lee et al., 1994a, 1994b; Marzoli et al. 1999), determined on volcanic massifs and lava flows from CVL and AP, do not clearly support a hot-spot origin as the model of Wilson-Morgan (Wilson 1965, 1973; Morgan 1971, 1972a,b). Then, the lack of age progression along the CVL magmatism renders the age determination of each volcano of the line indispensable to modelize the entire CVL. Moreover, the AP geochronology (Gouhier et al. 1974; Temdjim et al. 2004; Nkouandou et al. 2008 ; Fagny et al. 2012) evidenced that these volcanoes are relatively more recent (11 Ma to recent) than those of the CVL (up to 70 Ma to recent, Mbowou et al. 2012). Thus, whether the AP magmatism belongs to CVL, as proposed in the Y-shaped model for the whole CVL (Fitton, 1980,

1983), though largely debated (Gouhier et al. 1974; Aka et al. 2004, 2009; Nkouandou et al. 2010), remains questionable and requires many more age determinations. We contribute to this debate by producing a new geochronology base on the K-Ar dating of alkali intraplate volcanic massif of Tchabal Mbabo in the Cameroon volcanic Line.



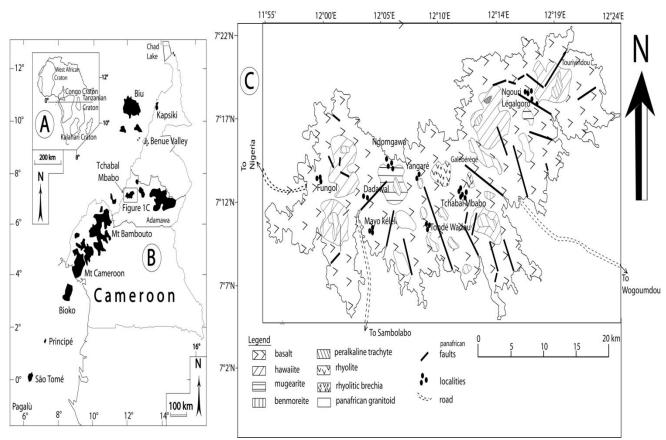


Fig. 1: A: Main Cratonsof Africa Plate (after Kampunzu&Popoff1991), B: Cameroon Volcanic Line (after DéruelleEt Al. 1991, Modified) and C: the Geological Map of Tchabal Mbabo Volcano.

## 2. Geological setting

The Cameroon Volcanic Line (CVL), in central Africa (Fig. 1), is a 1600 km long volcanic chain which straddles both the Atlantic Ocean crust and continental Pan-African crust of Cameroon, from Pagalu in the Gulf of Guinea to nearby Lake Chad (Déruelle et al. 2007). It intersects the Adamawa plateau at the Tchabal Mbabo volcano (Déruelle et al. 1990) and ends in the Chad Lake neighbourhood, in the northernmost Cameroon and southern Chad Republic. To explain the formation of the CVL, numerous interpretations were discussed, the most important of which can be found in Déruelle et al. (1991). The most widely accepted structural model is that CVL is due to a rejuvenation of the pan-African fracture zone, N70°E oriented, at the beginning of the Central Atlantic Ocean opening (Moreau et al. 1987). The Tchabal Mbabo basement consists of granitic and gnessic rocks, rising up to 2100 m. Those rocks are probably Pan African in age, as shown by those described at Méiganga (Ganwa et al. 2008) and northern Ngaoundéré, with ages of 610-652±54 Ma and 926±12 Ma (Tchameni et al. 2006), which place them within the Central Pan-African Chain (CPAC). CVL is composed of volcanic centres and anorogenic complexes with volcanic rocks, forming a chain of Tertiary to recent lavas (Déruelle et al. 1991; Kamden et al. 2002; Kamgang et al. 2010; Mbowou et al. 2012). Continental magmas evolve towards peralkaline rhyolite lavas, whereas those in the oceanic sector evolve towards phonolites (Fitton 1987). This points out the progressive crustal contamination of the continental magmas, accompanied by a crystal fractionation, to explain this distinction.

## 3. Analytical methods

Petrographic analyses of the studied rocks were carried out on 7 thin sections, prepared from the most representative samples of the lavas at the Laboratory of Geosciences GEOPS, University Paris-Sud Orsay and Centre de Recherches Pétrographiques et Géochimiques (CRPG) of Nancy, both in France. Major element analyses have been carried out by ICP-OES (Inductively Coupled Plasma - Optical Emission Spectrometry) and trace element analyses by ICP-MS (Inductively Coupled Plasma - Mass Spectrometry). During the preparation of samples, about 300 mg of powder have been treated with LiBO<sub>2</sub> and dissolved in HNO<sub>3</sub>. Five international geostandards have been used, as follows: basalt BR, diorite DRN, serpentinite UBN, anorthosite ANG and granite GH (Carignan et al. 2001).

Minerals have been analysed with a CAMECA SX 100 electron microprobe (15 kV, 10 nA), at Université Pierre et Marie Curie, Paris VI, France; for this purpose, K $\alpha$  lines were used. The chemical composition was determined using the following standards: diopside for Si, Ca and Mg; Fe<sub>2</sub>O<sub>3</sub> for Fe; MnTiO<sub>3</sub> for Ti and Mn; Cr<sub>2</sub>O<sub>3</sub> for Cr; albite for Na; orthoclase for K and Al. The counting times were 10 s for Pk (peaks) and 10 s for Bg (background), with a 5 µm defocused beam. As standards, a combination of natural and synthetic minerals were used. Data corrections were made using the PAP method correction of Pouchou and Pichoir (1991).

Three samples (M4, M12 and MB17) were K-Ar dated, at the Laboratoire des Sciences du Climat et Environnement, Gif-Sur-Yvette, France.

Groundmass was separated from phenocrysts and other phases using a Sm-Co magnetic and heavy liquid, followed by handpicking under a binocular microscope.

All the samples were dated using the unspiked K-Ar technique described by Charbit et al. (1998). This technique differs from the conventional isotope dilution method in that argon extracted from the sample is measured in sequence with purified aliquots of atmospheric argon at the same working gas pressure in the mass-spectrometer. The determination of K was carried out by atomic absorption with a relative precision of 1%. Argon was extracted by radio frequency heating of a 300 mg sample in an ultra-high-vacuum glass line and purified with Ti sponge and Zr-Al getters. Isotopic analyses were performed on total <sup>40</sup>Ar contents ranging between 2.9  $10^{-11}$  and 1.0 x  $10^{-10}$ mol using a 180°, 6 cm radius mass spectrometer equipped with a double faraday collector allowing the simultaneous measurement of <sup>36</sup>Ar and <sup>40</sup>Ar at an ac-

celerating potential of 620 V. The manometric calibration is based on periodic replicate determinations of international dating standards of known K-Ar age using the same procedure as for the unknown samples to be measured as described in Charbit et al. (1998). This allows the total <sup>40</sup>Ar content of the sample to be determined with a precision of about  $\pm 0.2\%$  (2 $\sigma$ ). Used standard is HD-B1 dated 24.21  $\pm 0.32$  Ma, (Fuhrmann et al. 1987; Hess and Lippolt 1994; Hautmann and Lippolt 2000).

### 4. Results

#### 4.1. K-Ar geochronology

Lavas have been identified according to IUGS classification (after Le Maître, 2002) as: basalt (M4) (45 wt% < SiO<sub>2</sub>< 52 wt% and normative olivine of 8 %); basanite (MB12) (SiO<sub>2</sub>< 45 wt% and

normative olivine > 10 wt %); and trachyandesite (57 wt % < SiO\_2< 63 wt %, MB17).

The K-Ar ages are reported in Table 1. For each sample, three independent measurements of K were carried out. Potassium concentrations were combined to yield a mean value. Age determinations of each sample were made using this value and the weighted mean of the two independent measurements of <sup>40</sup>Ar\*(radiogenic argon). Uncertainties for the Ar data are 1 $\sigma$  analytical only, and consist of propagated and quadratically averaged experimental uncertainties arising from the <sup>40</sup>Ar (total), and <sup>40</sup>Ar\* determinations. Uncertainties on the ages are given at  $2\sigma$ .

Ages are  $28.88 \pm 0.61$  Ma (basalt),  $28.60 \pm 0.60$  Ma (basanite) and  $38.22 \pm 0.80$  (trachyandesite) (Table 1). Therefore, the Tchabal Mbabo volcanic massif was emplaced in his present area during Paleogene, i.e., end Eocene-Oligocene.

Table 1: K-ArAge Data of Lavas									
Sample	K (wt %) $\pm 2\sigma$	Mass molten (g)	<sup>40</sup> Ar %	${}^{40}Ar \% \ 10^{-11} \\ (Mol./g) \pm 1\sigma$	$^{40}$ Ar % 10 <sup>-11</sup> (Mol./g) ± 1 $\sigma$ Weighted mean	Age (Ma) ±2σ			
M12									
8711	$1.544 \pm 0.015$	0.31658	83.216	$7.681 \pm 0.039$					
8726	« »	0.31879	84.052	$7.755 \pm 0.039$	$7.718 \pm 0.027$	$28.60 \pm 0.60$			
M4									
8724	$1.004 \pm 0.014$	0.30405	22.190	$5.073 \pm 0.025$					
8725	« »	0.34671	23.361	$5.065 \pm 0.025$	$5.069 \pm 0.018$	$28.88 \pm 0.61$			
MB17									
8712	$4.317 \pm 0.043$	0.30645	89.421	$28.710 \pm 0.144$					
8727	« »	0.31432	70.408	$29.132 \pm 0.146$	28.918 ±0.102	$38.22 \pm 0.80$			

#### 4.2. Petrography and mineralogy of dated lavas

Two dated basalt-basanite lavas have microlitic porphyritic texture (Fig. 2. A, B, C, D, E and F). They are composed of olivine, clinopyroxene and oxides phenocrysts, disposed in a microlitic matrix consisting of the same minerals as the phenocrysts, plus plagioclases. Brown crystals of amphibole are present in some samples (Fig. 2C). Chemical compositions of representative minerals of dated lavas are presented in Table 2.

Olivine phenocrysts (size of 250  $\mu$  to 1,2 mm), containing rare oxide microcrysts (Fig. 2A), are less abundant (7 to 10 modal%); they are euhedral and Mg-rich (Fo<sub>64,18-65.76</sub>). Olivine microcrysts are less magnesian (Fo<sub>54</sub>). CaO contents are relatively higher in the core (0.34-0.41 wt %) compared to rims (0.33 wt%); NiO contents are low (0.1 wt%) in phenocrysts.

Clinopyroxene crystals (60  $\mu$ m to 2mm) are euhedral (Fig. 2B) and have an abundance of 10 modal %. Some phenocrysts exhibiting corroded rimes are subhedral. Enclosed oxide microcrysts are frequent. All clinopyroxenes show a diopside (Wo<sub>49,4-49,7</sub>En<sub>40,4-40,5</sub>Fs<sub>9,9-9,1</sub>) to augite (Wo<sub>42,0-43,3</sub>En<sub>46,2-46,9</sub>Fs<sub>9,8-11,8</sub>) composition (classification scheme of Morimoto et al. 1988). M12 clinopyroxene crystals are richer in  $TiO_2$  (1.8-2.2 wt %) and  $Al_2O_3$  (5 wt%), relative to M4 clinopyroxene ( $TiO_2$ : 0.2-0.3 wt% and  $Al_2O_3$ : 2.0-2.6 wt%).

The plagioclase microlites are dominant within the matrix (60 in volume %). They have acicular shapes and a labradorite composition ( $An_{60.64-58.28}Ab_{37.14-38.95}$ ).

The oxides are mainly represented by Ti-magnetite, showing a  $TiO_2$  amount of 15-19 wt% and a FeO abundance ranging between 77.59 and 78.55 wt%.

Amphiboles appear as euhedral phenocrysts, always crowned by a thin layer consisting of oxide microcrysts. All crystals have a pargasite composition, as shown by the Si abundance (0.658-0.676 apuf) and Mg/ (Mg+Fe<sup>2+</sup>) ratio (64.66-68.59) (Leake et al. 1998).

Trachyandesite lava exhibits a microlitic porphyric texture. It is composed of phenocrysts of ferroaugite ( $Wo_{43.54-44.66}En_{14.62-11.91}Fs_{45.71-43.44}$ ) and microlites of andesine ( $Or_{7.38-9.69}Ab_{62.07-59.46}An_{30.55-30.84}$ ), pargasite, alkali feldspar, oxides ( $TiO_2$ : 14-18 wt%,  $Al_2O_3$ : 77-78 wt %), and apatite. The groundmass contains microlites of oligoclase ( $Or_{23.07-11.46}Ab_{64.26-58.55}An_{12.67-29.99}$ ) and also of oxides, ferroaugite, feldspar and apatite.

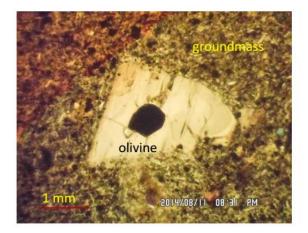


Fig. 2A. Corroded rim of olivine crystal enclosing oxide crystal in the groundmass of plagioclase and oxide microlits of basanite M4.

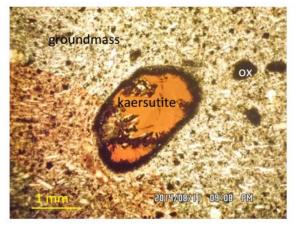


Fig. 2C. Kaersutite amphibole phenocryst encircled by a crown of oxide microcrysts (LAP). Sample M4

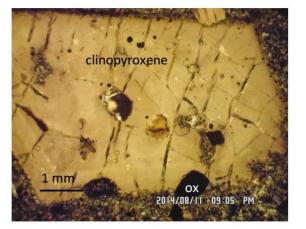


Fig. 2B. Euhedral shape of zoned clinopyroxene crystal in the groundmass of microcrystalized plagioclase and oxides. Basanite M12.

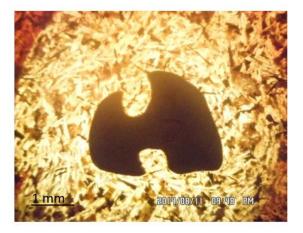
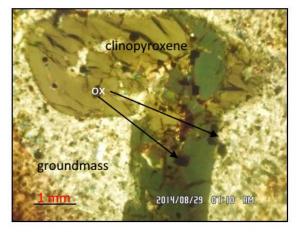


Fig. 2D. Corroded xenocryst of spinel in basanite M12 (LAP).



**Fig. 2E.** Aggregation of skeletal green clinopyroxene phenocrysts, containing angular oxides crystals (LPA). Sample MB17

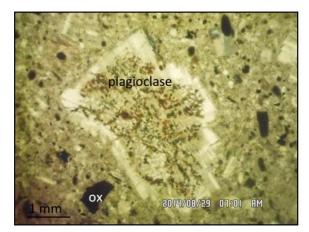


Fig.2F. Resorbed plagioclase phenocrysts, containing tiny oxide crystals (LPA). Sample MB17

Fig. 2: Microphotographs of Studied Lavas.

rock	basa- nite			Table	e 2: Che	emical (	Compo	sition o	f Repre	sentati	ve Mine	erals of	f Dated	Lavas trachy- andesite					
mineral	ol				срх				pl					ox			par		
sample	M4				M12		M4		M12					MB17			g M1 2		
description	ph.c			mic	ph.c	ph.c	ph.r	ph.c	ph.c	ml	ph.c	ph. b	ph.c	ph.c		mic	2 ph.c	ph.r	ph.c
SiO <sub>2</sub> (wt%)	37.39	37.3	37.2	35.9	46.2	45.1	50.	51.1	52.7	52.	53.1	55.	55.	0.06	0.0	0.1	39.	39.	40.
TiO <sub>2</sub>		2	7	2	0 1.77	7 2.19	85 0.2	2 0.30	0	73	7	13	47	15.17	5 14.	3 18.	57 2.6	54 2.7	59 2.4
Al <sub>2</sub> O <sub>3</sub>					5.17	5.14	4 1.9	2.56	29.1	28.	28.9	27.	27.	0.64	77 0.7	98 0.5	6 13.	7 13.	6 13.
FeOt	31.67	30.1	31.3	38.3	11.7	12.1	8 10.	9.99	7 0.81	81 0.7	6 0.76	60 0.3	12 0.4	78.55	4 78.	6 77.	65 11.	85 11.	62 10.
MnO	0.49	5 0.52	0 0.53	1 0.75	8 0.21	2 0.20	93 0.2	0.23	0.01	1	0170	6	6	0.92	52 0.9	59 0.9	73 0.1	36 0.1	41 0.0
MgO	31.84	32.4	31.7	25.1	10.6	10.2	4 14.	14.6						0.15	6 0.1	9 0.0	6 12.	6 12.	6 12.
-		9	7	0	6 22.6	6 22.1	39 20.	2 21.4	12.6	12.	12.2	10.	10.	0.15	4	1	21 12.	18 12.	83 12.
CaO NiO	0.34 0.13	0.35 0.02	0.41 0.09	0.33 0.00	5	5	49	5	1	28	6	45	12				20	42	30
Na <sub>2</sub> O					0.61	0.78	0.3 7	0.31	4.27	4.5 8	4.50	5.4 5	5.5 8				2.5 2	2.4 0	2.5 1
K <sub>2</sub> O									0.39	0.4 1	0.43	0.4 9	0.6 5				1.2 0	1.2 4	1.1 8
Total	101.8 8	100. 84	101. 38	100. 41	99.0 55	98.0 22	99. 13	100. 27	100. 05	99. 62	100. 14	99. 56	99. 39	95.49	95. 19	98. 28	95. 90	95. 93	95. 96
Si (apuf)	0.999	$1.00 \\ 2$	1.00 0	1.01 3	1.76 8	1.75 1	1.9 09	1.89 3	2.39 6	2.4 08	2.41 3	2.5 01	2.5 19	0.002	0.0 02	0.0 05	0.6 59	0.6 58	0.6 76
Ti		2	0	5	0.05 1	0.06 4	0.0 07	0.00 8	0	00	5	01	1)	0.451	0.4 40	0.5 47	0.0 33	0.0 35	0.0 31
Al					0.23 3	0.23 5	0.0 87	0.11 2	1.56 4	1.5 50	1.54 9	1.4 76	1.4 52	0.030	0.0 35	0.0 25	0.2 68	0.2 72	0.2 67
Fe++	0.708	0.67 7	0.70 3	0.90	0.18	0.17	0.2	0.19	4	50	2	70	52	1.415	1.4	1.5	0.1	0.1	0.1
Fe+++		/	3	4	4 0.17	8 0.19	25 0.1	1 0.10	0.03	0.0	0.02	0.0	0.0	1.062	03 1.0	21 0.8	63	58	45
Mn	0.011	0.01	0.01	0.01	4 0.00	3 0.00	06 0.0	7 0.00	1	27	9	14	17	0.031	79 0.0	68 0.0	0.0	0.0	0.0
Mg	1.268	2 1.30	2 1.27	8 1.05	7 0.60	6 0.59	$\begin{array}{c} 08 \\ 0.8 \end{array}$	7 0.80						0.009	32 0.0	32 0.0	02 0.3	02 0.3	01 0.3
Ca	0.010	0 0.01	1 0.01	5 0.01	8 0.92	3 0.92	$\begin{array}{c} 06 \\ 0.8 \end{array}$	7 0.85	0.61	0.6	0.59	0.5	0.4	01009	08	00	03 0.2	02 0.2	18 0.2
Ni	0.003	0 0.00	2 0.00	$\begin{array}{c} 0 \\ 0.00 \end{array}$	9	0	24	1	4	01	6	08	93				18	22	19
	0.003	0	2	0	0.04	0.05	0.0	0.02	0.37	0.4	0.39	0.4	0.4				0.0	0.0	0.0
Na					5	9	27	3	6 0.02	06 0.0	6 0.02	80 0.0	91 0.0				81 0.0	77 0.0	81 0.0
K					49.3	49.7	42.	43.2	2	24	5	29	38				25	26	25
Wo					9 40.5	1 40.4	03 46.	9 46.8											
En					3 10.0	3	19 11.	6											
Fs Mg/Mg+Fet		65.7	64.4	53.8	8 76.8	9.86 76.8	78 78.	9.85 80.8									64.	65.	68.
+Mn	64.18	6	0	6	2	9	15	8		22		20	2.6				66	31	59
Or (%)									2.22	2.3 4 20	2.47	2.8 1 47	3.6 7						
Ab									37.1 4	39. 37	38.9 5	47. 22	48. 11						
An									60.6 4	58. 28	58.5 8	49. 97	48. 22		<i>.</i> .	0 -			
%Usp														0.46	0.4 5	0.5 6			

ol: olivine, cpx: clinopyroxene, pl: plagioclase, ox: oxides, parg: pargasite, ph.c: phenocryst core, ph.r: phenocryst rime, mic: microcryst, ml: microlite.

#### 4.3. Geochemistry of dated lavas

The abundance of major and trace elements of selected lavas used in this study are presented in Table 3. Major elements exhibit features typical to alkali lavas worldwide, such as high  $TiO_2$  (up to 4wt%) and high sum of alkalis (between 4 and 10 wt%).

Contents of REE within the lavas (42 to 124 times the mantle values) are normalized to Primitive Mantle (after Hofmann 1988)

and illustrated in Fig. 3. All lavas are characterized by the very high contents of LREE relative to HREE, exhibiting a gentle slope pattern. There is no particular anomaly, even in Eu, while CeN/YbN ratios of basaltic lavas range between 7.5 and 12.1.

Trace elements contents of studied lavas also were normalized to Primitive Mantle (after Hofmann 1988). These patterns, presented in Fig. 4, are similar to those of Cenozoic alkali lavas studied in the Cameroon Line (Kamgang et al. 2010), Adamawa plateau (Nono et al. 1994; Nkouandou et al. 2008 & 2010) and world-wide. The patterns display roughly the same configuration, with increased contents showed by Rb and especially Ba, and followed by regularly decrease towards Lu. Negative anomalies are recorded by Th, U, K, P and, partially, Ti. Positive anomalies are noticed for Nb, Ta

and Sr. For trachandesite, there are noteworthy negative anomalies in P, Sr and Ti. This pattern is particular in positive K anomaly. Ratios of Nb/Ta (13.3-1.9) and Zr/Hf (39.2-43.8) are roughly constant.

Table 3: Whole Rock	Chemical .	Analyses	of Studied Lavas

Lava		le Rock Chemical Analyses of		
Lava	Basalt	Basanite	Trachyandesite	
Sample	M4	M12	MB17	
(wt%)				
SiO <sub>2</sub>	46.41	42.97	60.35	
TiO <sub>2</sub>	3.25	3.99	0.60	
$Al_2O_3$	15.83	14.94	14.75	
Fe <sub>2</sub> O <sub>3</sub>	11.93	12.70	9.04	
MnO	0.17	0.21	0.23	
MgO	5.95	6.35	0.25	
CaO	9.68	10.02	2.67	
Na <sub>2</sub> O	2.83	3.90	4.54	
K <sub>2</sub> O	1.09	1.64	5.21	
$P_2O_5$	0.39	0.96	0.16	
LOI	3.14	1.37	1.44	
sum	100.67	99.07	99.22	
CIPW norm (%)				
	0.00	0.00	5.75	
Quartz	0.00	0.00	5.65	
Corundum	0.00	0.00	0.00	
Orthoclase	6.44	9.70	30.78	
Albite	23.98	12.69	38.43	
Anorthite	27.26	18.43	4.46	
Leucite	0.00	0.00	0.00	
Nepheline	0.00	10.99	0.00	
Diopside	14.91	20.46	6.82	
Wollastonite	0.00	0.00	0.00	
Hypersthene	4.17	0.00	6.62	
Olivine	10.59	12.30	0.00	
Magnetite	2.06	2.19	2.78	
Ilmenite	6.17	7.58	1.14	
Haematite	0.00	0.00	0.00	
Apatite	0.90	2.22	0.37	
(ppm)				
Be	0.91	1.70	2.62	
Rb	33.12	36.95	79.5	
Sr	827	1160	102	
Cs	0.64	0.42	0.22	
Ba	371	690	1545	
V	253.7	263.6	<d.l.< td=""><td></td></d.l.<>	
Cr	130	147	6	
Co	41.97		0.96	
		37.47		
Ni	54.53	55.29	<d.1.< td=""><td></td></d.1.<>	
Cu	31.60	39.22	14.17	
Zn	130	133	163	
Y	23.94	31.8	44.74	
Zr	190	280	607	
Nb	33.68	76.24	69.45	
Hf	4.85	6.38	14.16	
Та	2.53	5.54	4.99	
Th	2.41	4.25	8.70	
U	0.66	1.23	2.14	
Pb	9.15	2.943	28.90	
La	25.87	55.42	76.28	
Ce	58.67	115.90	157.90	
Pr	7.59	14.20	18.75	
Nd	31.81	57.44	73.34	
Sm	6.92	11.31	14.22	
Eu	2.34	3.67	5.44	
Gd	5.97	9.10	11.25	
Tb	0.86	1.25	1.61	
Dy	4.95	6.84	9.21	
Но	0.94	1.27	1.76	
Er				
	2.31	3.05	4.53	
Tm	0.31	0.39	0.63	
Yb	2.01	2.49	4.28	
Lu	0.29	0.35	0.67	
Sample location: M4: N07 25463	2º E012 202459 M12 N072506	00E010201770, MD 17, NO7 205	45 E012 259609	

Sample location: M4: N07.25463°, E012.29345°; M12: N0725969°E01230177°; MB17: N07.29545 E012.25869°.

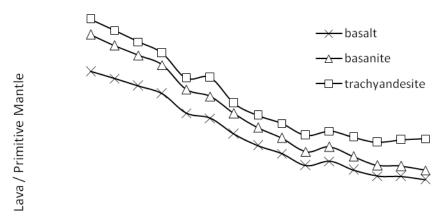


Fig. 3:REE Patterns of Dated Lavas from Tchabal Mbabo (Normalization after Hofmann 1988).

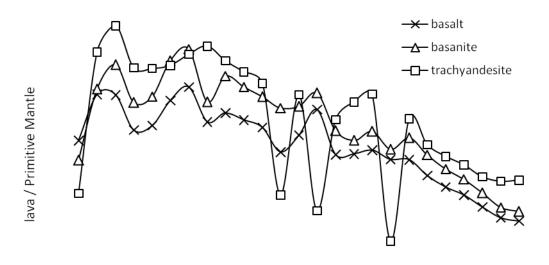


Fig. 4: Primitive Mantle Normalized Multi-Element Diagrams for Dated Lavas from Tchabal Mbabo Volcano (Normalization after Hofmann 1988).

# 5. Discussions

The K-Ar ages of 28.88±0.61 Ma, 28.60±0.60 Ma and 38.22±0.80 Ma suggest that more than one eruptive episode have occurred at Tchabal Mbabo volcanic massif. The differentiate trachyandesite is 10 Ma older than the basaltic lavas. This indicates that basaltic and trachyandesite do not belong to the same lava series, although they may evolve from the same source or sources of similar composition. At more than 200 km north and east from Ngaoundéré, K-Ar ages of 7.8±1.4 Ma have been obtained for a basalt sample, as well as 8.3±0.5 - 10.9±0.4 Ma ages for four felsic lavas (Nkouandou et al. 2008). One basalt of ankaramitic composition and two trachytic lavas have also been dated as 7.2, 7.8 and 9.8 Ma at Tchabal Nganha volcano massif, by Gouhier et al. (1974). Moreover, Temdjim et al. (2004) determine the ages of some lavas from the Ngaoundéré area, as follows: 6.2±0.2 Ma for one hawaiite sample,  $6.5 - 6.8 \pm 0.1$  Ma for two phonolites and 0.91±0.06 Ma for a mugearite.

#### 5.1. Origin of the Cameroon volcanic line

It appears that the volcanic eruptions occurring in Adamawa plateau were more recent compared to Tchabal Mbabo volcanic activity and some of the major volcanic eruptions of the Cameroon Line (Fig. 5). At Mont Cameroon, the only active volcano massif of the Cameroon Line, K-Ar ages range from 2.83±0.11 to 0.00±0.09 Ma and from 20±3 to 29±2 Ma (Wandji et al. 2009). Therefore, it seems like Tchabal Mbabo volcano massif belongs rather to the Cameroon Volcanic Line than to Adamawa plateau province. However, a distinction between the Cameroon Volcanic Line and Adamawa volcanism is difficult to be made (Fitton 1983). Those observations show that the volcanic manifestations occur diversely along the Cameroon Line and Adamawa plateau, as shown in Fig. 6 (Déruelle et al. 2007, modified and references herein), probably according to the weaknesses conffered by the underneath continental crust in each volcanic area. In this regard, the formation of the Cameroon Line is difficult to council with

any space-time relationships, especially the hot-spot model of Wilson-Morgan (Wilson 1965, 1973; Morgan 1971, 1972a,b, 1981). The most realistic conception regarding the Cameroon line and Adamawa volcanism is that they have resulted from the rejuvenation of the Pan-African faults, inherited from the Pan-African orogeny and reactivated during Mesozoic and Cenozoic (Moreau et al. 1987). Those faults have crosscut the continental crust down to the mantle (Dorbath et al. 1984) and served as pathway for magma, to reach the surface. Thus, the hot-spot model suggested by Marzoli et al. (1999), also for the oceanic sector of the Cameroon Line volcanism, seems not to be realistic.

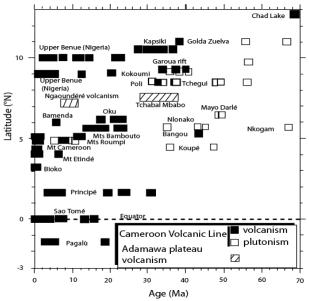
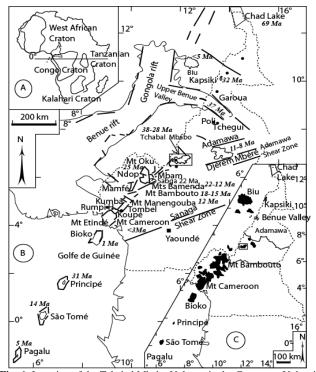


Fig. 5:Age Distribution of volcanoes and anorogenic plutons along the Cameroon Line (Déruelle et Al. 2007, Modified) and Adamawa Plateau (Nkouandou et al. 2008).



**Fig. 6:** Location of the Tchabal Mbabo Volcano in the Cameroon Volcanic Line Volcanism and Adamawa Plateau. (A): Main African Cartons (after Kampunzu&Popoff1991); (B): Horst Structure of the Cameroon Line (Modified); and (C): Mains Volcanoes Location along the Cameroon Line (after Déruelle et al. 1991).

#### 5.2. Origin of lavas

The origin of Tchabal Mbabo lavas is widely discussed by Fagny (2016). Determination of the partial melting rate for Tchabal Mbabo volcanism using the method of fractional melting has been attempted. The high Cen/Ybn ratios (7-12) are consisting with a low partial melting rate. A model of partial melting of about 2% of a deep mantle source (more than 80 km), containing garnet and small amounts of phlogopite is proposed for Tchabal Mbabo basanite lavas. The presence of garnet and phlogopite phases within the source is proved by the low values of heavy REE (Wilson 1973) and the negative K anomaly of those lavas (Chauvel et al. 1992). The magma mixing is presumed for basanite of Tchabal Mbabo as the occurrence of plagioclase and spinel xenocrysts in those lavas suggest (Fig. 2D). The pressures in the melting zone calculated using Scarrow and Cox (1995) formula are between 2.49 and 4.3 GPa, corresponding to depths of 82 and 143 km. These depths are consisting with a melting source located just under the sub-lithosphere mantle. The same hypothesis has previously been proposed by Nono et al. (1994) for the Nganha volcano in the Adamawa plateau and the Upper Benue valley volcanism (Ngounouno et al. 2003) in the Cameroon Volcanic Line.

As previously suggested, trachyandesite and basanite lavas do not belong to the same lava series. The pretrogenesis of trachyandesite lava is discussed in term of fractional crystallization process. The mineral phases involved in this process are apatite, alkali feldspar and Fe-Ti oxides. This hypothesis is substantiated by the P, Sr and Ti negative anomalies in the spider diagram (Fig. 4) while the positive anomalies in Nb and Ta exclude any interaction between trachyandesite lava and the continental crust materials.

# 6. Conclusions

Tchabal Mbabo volcano in the Adamawa plateau was built up during Eocene-Oligocene. Trachyandesites excepted, basalt and basanite lavas belong to the same magmatic series. Basanite lavas are the result of low partial melting of about 2% rate of the sublithospheric mantle source, containing residual garnet and phlogopite phases. The formation of the Cameroon Line is difficult to council with any space-time relationships, especially the hot-spot model suggested by the previous works.

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